

# The First Edition of the African Conference on High Energy Physics (ACHEP 2023) **Electromagnetic Effects in Deep Inelastic Neutrino-Proton Scattering** Konstantin Kouzakov<sup>1,2</sup>, Elena Kovalevskaia<sup>1</sup>, Alexander Studenikin<sup>1,2</sup>

## Introduction

Deep inelastic scattering on nucleons is the main process of interaction of h trinos with matter. The corresponding cross sections are measured, for IceCube experiment on the basis of the dependence of the neutrino flux range 60 TeV - 10 PeV on the direction of its arrival in the detector [1]. unique opportunity to study the effects of physics beyond the Standard Mo ultrahigh neutrino energies. In particular, the search for the electromagne neutrinos is of interest [2]. We have taken into account the contribution of and magnetic form factors in the theoretical formalism of the process of dee trino scattering on the proton. We obtained expressions for the differentia sections, which can be used to analyze and interpret the experimental data neutrino scattering on nucleons in order to find the electromagnetic property

### Neutrino electromagnetic properties

The electromagnetic interactions of neutrinos are one of the unsolved fur lems in neutrino physics. They are actively studied theoretically and sear experiments. The neutrino electromagnetic properties are directly related to ementary particle physics. For example, they can be used to distinguish bet Majorana neutrinos. Since the neutrino electromagnetic properties are absent in the Standard Model with massless neutrinos, their discovery will open a window to new physics [2].

### Neutrino electromagnetic form factors

The electromagnetic properties of neutrinos are contained in the vertex function (see fig. 1), which is a matrix in spinor space. It can be decomposed in terms of linearly independent products of Dirac matrices and 4-vectors l and l'.



Figure 1: The effective vertex of the neutrino electromagnetic interaction The electromagnetic vertex can be written as follows:

$$\Lambda_{\mu}(q) = f_Q\left(q^2\right)\gamma_{\mu} - f_M\left(q^2\right)i\sigma_{\mu\nu}q^{\nu} + f_E\left(q^2\right)\sigma_{\mu\nu}q^{\nu}\gamma_5 + +f_A\left(q^2\right)\left(q^2\gamma_{\mu} - q_{\mu}\not{q}\right)\gamma_5.$$
(1)

For brevity, we refer to the combination

 $f_Q - q^2 f_A$ as the *charge form factor*, desnoting it by  $f_O$ , and to the combination

$$f_M - i f_E$$

as the *magnetic form factor*, denoting it by  $f_M$ .

### 2.2 Experimental constraints

The research into the influence of electromagnetic fields on neutrinos and their electromagnetic properties has been conducted since the early 1960s. Since then, many dedicated experimental and theoretical studies have been carried out. Below we briefly outline the strongest constraints on the neutrino electromagnetic properties obtained in those studies.

### 2.2.1 Neutrino millicharge

The strongest limit on the neutrino electric charge (millicharge) obtained in experiments on neutrino scattering on a target is [3]

$$e_{\nu} \le 1.5 \cdot 10^{-12} e_0.$$

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It follows from the analysis of the data of the GEMMA experiment [4] on the scattering of reactor neutrinos on electrons.

The strongest limit on the neutrino millicharge obtained from astrophysical data is

 $e_{\nu} \leq 1.3 \cdot 10^{-19} e_0.$ 

It is based on the study of the motion of neutrinos having a nonzero electric charge in magnetized neutron stars [5].

### 2.2.2 Neutrino magnetic moment

One of the strongest of the available laboratory limits on the neutrino magnetic moment was also obtained in the GEMMA experiment [4]:

 $\mu_{\nu} \le 2.9 \cdot 10^{-11} \mu_B,$ 

where  $\mu_B$  is the Bohr magneton. Astrophysical constraints on the neutrino magnetic moment were obtained in [6–8], based on observations of luminosity of globular star clusters:

$$\left(\sum_{i,j} |\mu_{ij}|^2\right)^{1/2} \le (2.2 - 2.6) + \frac{1}{2}$$

### General formalism 3



Figure 2: Deep inelastic neutrino scattering on a nucleon The differential cross section for deep inelastic lepton-nucleon scattering in a laboratory frame is given by (see, for instance, [9])

$$\frac{d^2 \sigma^{\ell N}}{dE' d\Omega} = \frac{\alpha^2}{m_N q^4} \frac{E'}{E} \sum_{\substack{j=\gamma,\gamma Z, Z}} \eta_j L_j^{\mu\nu} W_{\mu\nu}^j, \tag{2}$$

where q is the momentum transfer, E and E' are the initial and final lepton energies, and  $L_{i}^{\nu\mu}$  and  $W_{\nu\mu}^{j}$  are the leptonic and hadronic tensors.

The total cross section derives from the expression

$$\sigma^{\ell N} = \int_{Q_0^2}^{s} dQ^2 \int_{Q^2/s}^{1} dx \, \frac{d^2 \sigma^{\ell N}}{dx dQ^2},\tag{3}$$

with  $Q_0 = 1$  GeV [10].

3.1 Notations

 $\eta_{\gamma} = 1, \qquad \eta_{\gamma Z} = \left(\frac{G_F M_Z^2}{2\sqrt{2}\pi\alpha}\right) \left(\frac{Q^2}{Q^2}\right)$  $Q^2 = -q^2 \approx 4EE' \sin^2 \frac{\theta}{2}, \qquad s = (\ell + P)^2 \approx m_N^2 + \nu = \frac{(q \cdot P)}{m_N} = E - \nu$ 

 $10^{-12}\mu_B.$ 

$$\frac{Q^2}{Q^2 + M_Z^2} \right), \qquad \eta_Z = \eta_{\gamma Z}^2,$$
$$+ 2m_N E, \qquad x = \frac{Q^2}{2\nu m_N},$$
$$-E', \qquad y = \frac{(q \cdot P)}{(\ell \cdot P)} = \frac{\nu}{E}.$$

### 4 Cross section

For the differential cross section of the  $\nu_i p$ -scattering, which accounts for electromagnetic  $(\gamma)$  and weak neutral-current  $(Z^0)$  interactions depicted in fig. 2, we get

$$\begin{aligned} \frac{f^{2}\sigma^{\nu,p}}{lxdQ^{2}} &= \frac{4\pi\alpha^{2}}{xQ^{4}} \Biggl\{ \sum_{f} \left| \mathbf{f}_{Q}^{fi}(Q^{2}) \right|^{2} \left[ xy^{2}F_{1}^{\gamma}(x,Q^{2}) + \left( 1 - y - \frac{x^{2}y^{2}m_{p}^{2}}{Q^{2}} \right) F_{2}^{\gamma}(x,Q^{2}) \right] \\ &- \eta_{\gamma Z} \operatorname{Re} \left\{ \mathbf{f}_{Q}^{ii}(Q^{2}) \right\} \left[ xy^{2}F_{1}^{\gamma Z}(x,Q^{2}) + \left( 1 - y - \frac{x^{2}y^{2}m_{p}^{2}}{Q^{2}} \right) F_{2}^{\gamma Z}(x,Q^{2}) \right. \\ &\left. \pm \left( y - \frac{y^{2}}{2} \right) xF_{3}^{\gamma Z}(x,Q^{2}) \right] \\ &\left. + \eta_{Z} \left[ xy^{2}F_{1}^{Z}(x,Q^{2}) + \left( 1 - y - \frac{x^{2}y^{2}m_{p}^{2}}{Q^{2}} \right) F_{2}^{Z}(x,Q^{2}) \right. \\ &\left. \pm \left( y - \frac{y^{2}}{2} \right) xF_{3}^{Z}(x,Q^{2}) \right] \\ &\left. + \frac{Q^{2}}{4xm_{e}^{2}} \sum_{f} \left| \mathbf{f}_{M}^{fi}(Q^{2}) \right|^{2} \left[ -xy^{2}F_{1}^{\gamma}(x,Q^{2}) + \frac{1}{2}(2 - y)^{2}F_{2}^{\gamma}(x,Q^{2}) \right] \Biggr\}, \end{aligned}$$

$$(4)$$

$$\begin{split} \frac{\partial}{\partial z} &= \frac{4\pi\alpha^2}{xQ^4} \Biggl\{ \sum_f \left| f_Q^{fi}(Q^2) \right|^2 \left[ xy^2 F_1^{\gamma}(x,Q^2) + \left( 1 - y - \frac{x^2 y^2 m_p^2}{Q^2} \right) F_2^{\gamma}(x,Q^2) \right] \\ &- \eta_{\gamma Z} \operatorname{Re} \left\{ f_Q^{ii}(Q^2) \right\} \left[ xy^2 F_1^{\gamma Z}(x,Q^2) + \left( 1 - y - \frac{x^2 y^2 m_p^2}{Q^2} \right) F_2^{\gamma Z}(x,Q^2) \\ &\pm \left( y - \frac{y^2}{2} \right) x F_3^{\gamma Z}(x,Q^2) \right] \\ &+ \eta_Z \left[ xy^2 F_1^Z(x,Q^2) + \left( 1 - y - \frac{x^2 y^2 m_p^2}{Q^2} \right) F_2^Z(x,Q^2) \\ &\pm \left( y - \frac{y^2}{2} \right) x F_3^Z(x,Q^2) \right] \\ &+ \frac{Q^2}{4xm_e^2} \sum_f \left| f_M^{fi}(Q^2) \right|^2 \left[ -xy^2 F_1^{\gamma}(x,Q^2) + \frac{1}{2}(2 - y)^2 F_2^{\gamma}(x,Q^2) \right] \Biggr\}, \end{split}$$
(4)

where "+" ("-") refers to neutrinos (antineutrinos).  $f_{Q,M}^{fi}$  are the neutrino charge and magnetic form factors of diagonal (i = f) and transition ( $i \neq f$ ) types given in units of elementary charge  $e_0$  and Bohr magneton  $\mu_B$ , respectively.  $F_{1,2,3}^{\gamma,Z,\gamma Z}$  are dimensionless structure functions of the proton. In the quark-parton model, one can use the Callan-Gross relation  $F_2 = 2xF_1$ . The total cross section is calculated on the basis of Eqs. (3) and (4).

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